
Multifluid module in Dyablo & Shamrock

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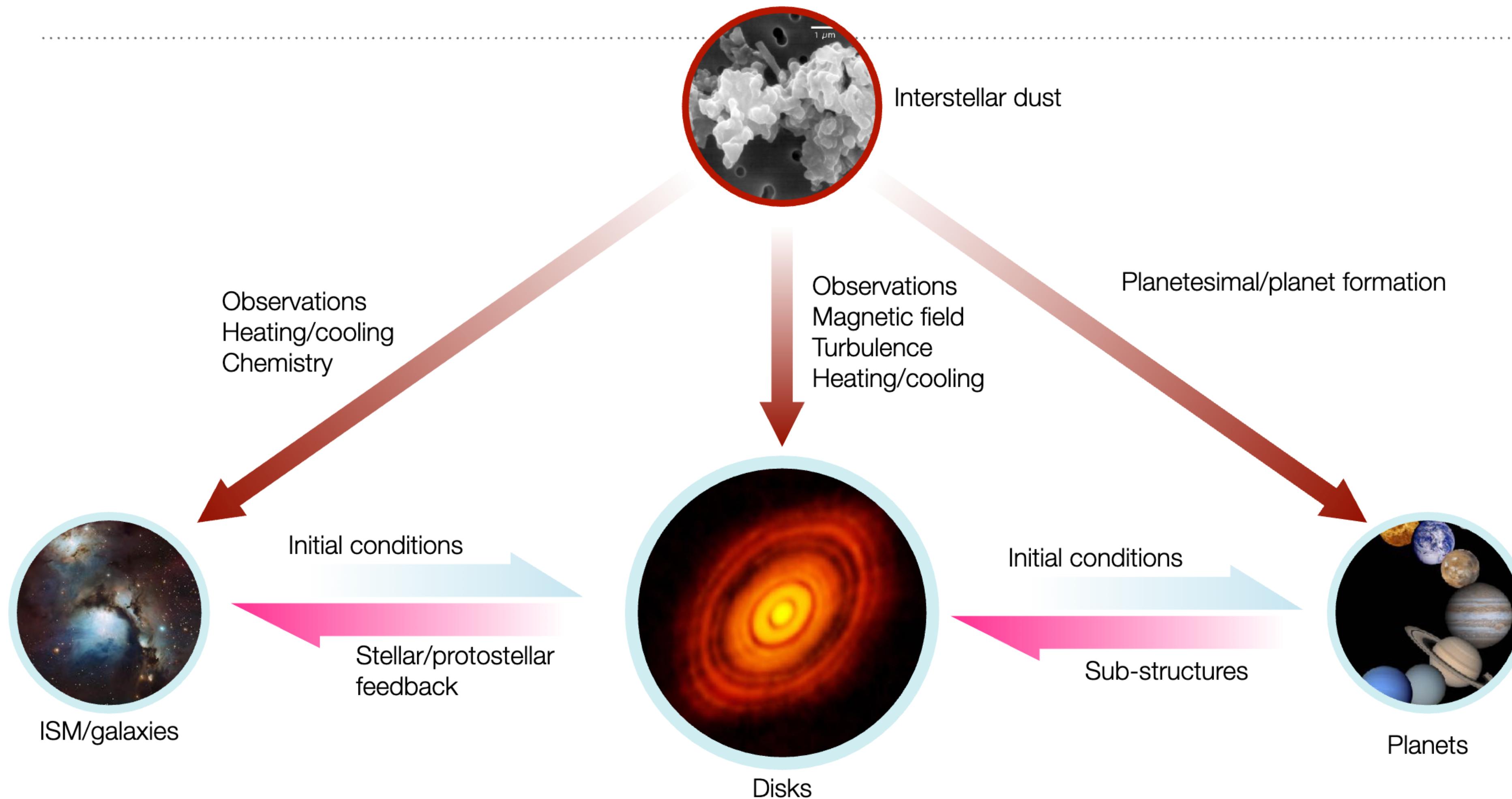


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Dust is everywhere

- Solid bodies made of silicates, carbon
- Only **1%** of the mass budget in the diffuse ISM ([Mathis et al. 1977](#); [Weingartner & Draine 2001](#))



Credit : Ugo Lebreuilly

Epstein drag :
(*Relevant for most
astrophysical object*)

$$F_{\text{drag}} \propto \frac{1}{T_s} ,$$

$$T_s \equiv \frac{\rho_{\text{grain}} s_{\text{grain}}}{\rho_g c_s} \sqrt{\frac{\pi \gamma}{8}}$$

Small grain / small stopping time

$\Rightarrow$

Strong coupling

Large grain / large stopping time

$\Rightarrow$

Weak coupling

$$\frac{\partial \rho_g}{\partial t} + \nabla \cdot (\rho_g \mathbf{v}_g) = 0,$$

$$\frac{\partial \rho_{d,i}}{\partial t} + \nabla \cdot (\rho_{d,i} \mathbf{v}_{d,i}) = 0,$$

$$\frac{\partial (\rho_g \mathbf{v}_g)}{\partial t} + \nabla \cdot (\rho_g \mathbf{v}_g \otimes \mathbf{v}_g + P_g \mathbf{I}) - \rho_g \mathbf{f}_g = \rho_d \frac{\mathbf{v}_d - \mathbf{v}_g}{T_s},$$

$$\frac{\partial (\rho_d \mathbf{v}_d)}{\partial t} + \nabla \cdot (\rho_d \mathbf{v}_d \otimes \mathbf{v}_d) - \rho_d \mathbf{f}_d = \rho_d \frac{\mathbf{v}_g - \mathbf{v}_d}{T_s}, \quad \text{pressureless dust fluid}$$

$$\frac{\partial E_g}{\partial t} + \nabla \cdot \left[(E_g + P_g) \mathbf{v}_g \right] - \rho_g \mathbf{f}_g \cdot \mathbf{v}_g = \rho_d \frac{\mathbf{v}_d - \mathbf{v}_g}{T_s} \cdot \left((1 - w) \mathbf{v}_g + w \rho_d \right)$$

$$\frac{\partial \mathbf{U}}{\partial t} + \nabla \cdot F(\mathbf{U}) + \mathbf{f}_{\text{src,ext}} = \mathbf{f}_{\text{drag}}$$

with $\mathbf{U} \equiv$

$$\begin{bmatrix} \rho_g \\ \rho_g \mathbf{v}_g \\ E_g \\ \rho_{d,1} \\ \rho_{d,1} \mathbf{v}_{d,1} \\ \vdots \\ \rho_{d,n} \\ \rho_{d,n} \mathbf{v}_{d,n} \end{bmatrix} F(\mathbf{U}) \equiv$$

$$\begin{bmatrix} \rho_g \mathbf{v}_g \\ \rho_g \mathbf{v}_g \otimes \mathbf{v}_g + P_g \mathbf{I} \\ (E_g + P_g) \mathbf{v}_g \\ \rho_{d,1} \mathbf{v}_{d,1} \\ \rho_{d,1} \mathbf{v}_{d,1} \otimes \mathbf{v}_{d,1} \\ \vdots \\ \rho_{d,n} \mathbf{v}_{d,n} \\ \rho_{d,n} \mathbf{v}_{d,n} \otimes \mathbf{v}_{d,n} \end{bmatrix} \mathbf{f}_{\text{src,ext}} \equiv$$

$$\begin{bmatrix} 0 \\ \rho_g \mathbf{f}_g \\ \rho_g \mathbf{f}_g \cdot \mathbf{v}_g \\ 0 \\ \rho_{d,1} \mathbf{f}_{d,1} \\ \vdots \\ 0 \\ \rho_{d,i} \mathbf{f}_{d,n} \end{bmatrix} \mathbf{f}_{\text{drag}} \equiv$$

$$\begin{bmatrix} 0 \\ \sum_{i=1}^N \rho_{d,i} \frac{\mathbf{v}_{d,i} - \mathbf{v}_g}{T_{s,i}} \\ \sum_{i=1}^N \rho_{d,i} \frac{\mathbf{v}_{d,i} - \mathbf{v}_g}{T_{s,i}} \cdot \mathbf{v}_g + w \sum_{i=1}^N \rho_{d,i} \frac{(\mathbf{v}_{d,i} - \mathbf{v}_g)^2}{T_{s,i}} \\ 0 \\ \rho_{d,1} \frac{\mathbf{v}_g - \mathbf{v}_{d,1}}{T_{s,1}} \\ \vdots \\ 0 \\ \rho_{d,n} \frac{\mathbf{v}_g - \mathbf{v}_{d,n}}{T_{s,n}} \end{bmatrix}$$

Numerical scheme

	0	1	2	...	$i - 1$	i	$i + 1$	...	$N - 3$	$N - 2$	$N - 1$	
	$\mathbf{U}_0^n$					$\mathbf{U}_i^n$					$\mathbf{U}_{N-1}^n$	

$$\frac{\partial \mathbf{U}}{\partial t} + \nabla \cdot \mathbf{F}(\mathbf{U}) + \mathbf{f}_{\text{src,ext}} = 0$$

	0	1	2	...	$i - 1$	i	$i + 1$	...	$N - 3$	$N - 2$	$N - 1$	
	$\tilde{\mathbf{U}}_0^n$					$\tilde{\mathbf{U}}_i^n$					$\tilde{\mathbf{U}}_{N-1}^n$	

$$\frac{d\tilde{\mathbf{U}}_0^n}{dt} = \Omega \tilde{\mathbf{U}}_0^n$$

$$\frac{d\tilde{\mathbf{U}}_i^n}{dt} = \Omega \tilde{\mathbf{U}}_i^n$$

$$\frac{d\tilde{\mathbf{U}}_{N-1}^n}{dt} = \Omega \tilde{\mathbf{U}}_{N-1}^n$$

	0	1	2	...	$i - 1$	i	$i + 1$	...	$N - 3$	$N - 2$	$N - 1$	
	$\mathbf{U}_0^{n+1}$					$\mathbf{U}_i^{n+1}$					$\mathbf{U}_{N-1}^{n+1}$	

Transport
Step

Drag
Step

Full time
step

$$\frac{d\tilde{\mathbf{U}}}{dt} = \mathbf{\Omega} \tilde{\mathbf{U}}, \quad \text{where} \quad \mathbf{\Omega} \equiv \begin{bmatrix} -\sum_{i=1}^N \epsilon_i \alpha_i & \alpha_1 & \alpha_2 & \dots & \alpha_N \\ \epsilon_1 \alpha_1 & -\alpha_1 & 0 & \dots & 0 \\ \epsilon_2 \alpha_2 & 0 & -\alpha_2 & \dots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ \epsilon_N \alpha_N & 0 & 0 & \dots & -\alpha_N \end{bmatrix} \quad \text{and} \quad \begin{aligned} \epsilon_i &\equiv \frac{\rho_{d,i}}{\rho_g} \text{ (dust-to-gas-ratio)} \\ \alpha_i &\equiv \frac{1}{T_{s,i}} \text{ (collision rate)} \end{aligned}$$

First-order Runge-Kutta implicit scheme in **FARGO3D** ([Benitez-Llambert et al., 2019](#)), **RAMSES** ([Verrier et al., 2025](#))

Second-order Van Leer, Second-order Runge-Kutta and other schemes in **ATHENA++** ([Huang & Bai, 2022](#))

Second-order Diagonal implicit Runge Kutta scheme ([Krapp et al., 2024](#))

$$\frac{d\tilde{\mathbf{U}}}{dt} = \Omega \tilde{\mathbf{U}} \quad \Rightarrow \quad \tilde{\mathbf{U}}(t) = e^{t\Omega} \tilde{\mathbf{U}}(t = t_0)$$

High accurate scheme

General (sparse/dense)

Minimal cost

GPUs friendly

$$e^A = \left(e^{A/n}\right)^n, \quad n = 2^s \in \mathbf{N}.$$

$$\tilde{A} \simeq e^{2^{-s}A}$$

Scaling step

$\Rightarrow$

$$e^A \simeq \tilde{A}^{2^s}$$

Squaring step

*Backward error bounded by the unit
roundoff*

Scaling factor *s* and Taylor polynomial
order *m* :

Minimal cost (scaling step)

avoid overshooting (squaring step)

Dyablo



IRK1, IRK2, MDIRK, EXPO

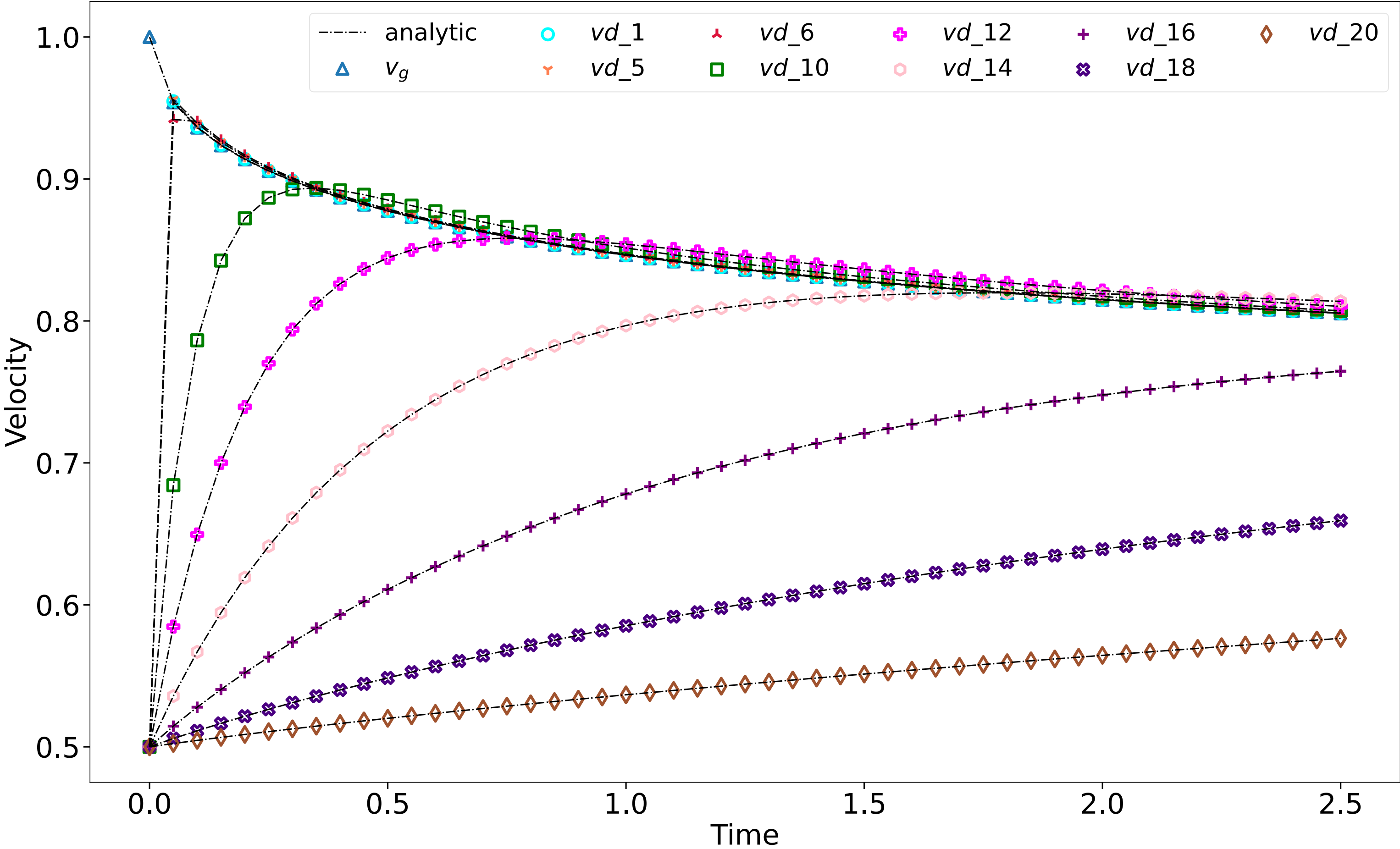
Shamrock



IRK1, EXPO

$$\frac{d\tilde{U}}{dt} = \Omega \tilde{U}$$

Dusty collision : 20 dust species



$$\frac{d\tilde{\mathbf{U}}}{dt} = \Omega \tilde{\mathbf{U}}$$

Error $\propto \Delta t$,

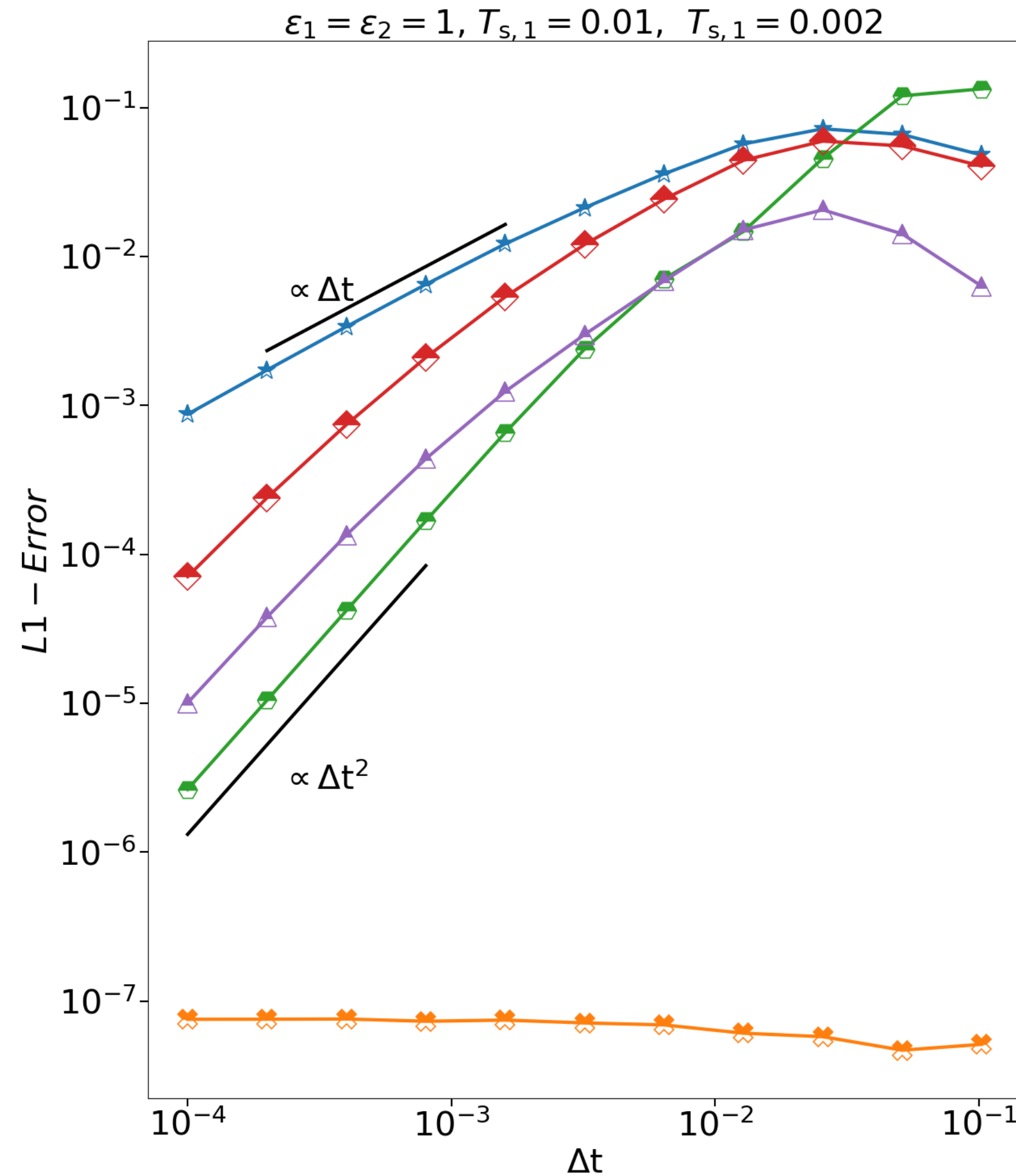
Time-to-solution $\propto 1/\Delta t$

10^4 (2 dust species)

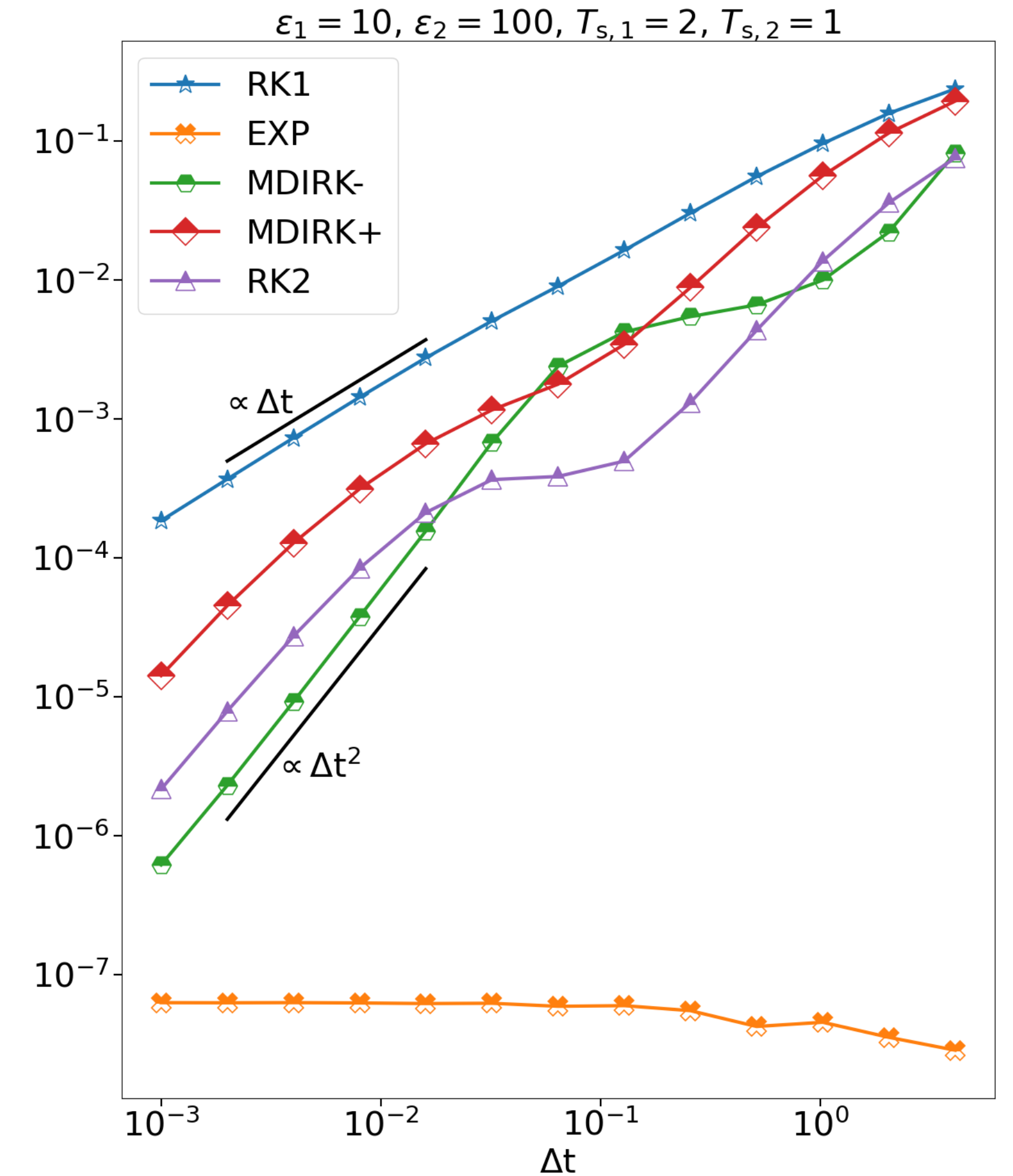
55 (20 dust species)

Sewanou et al, 2025

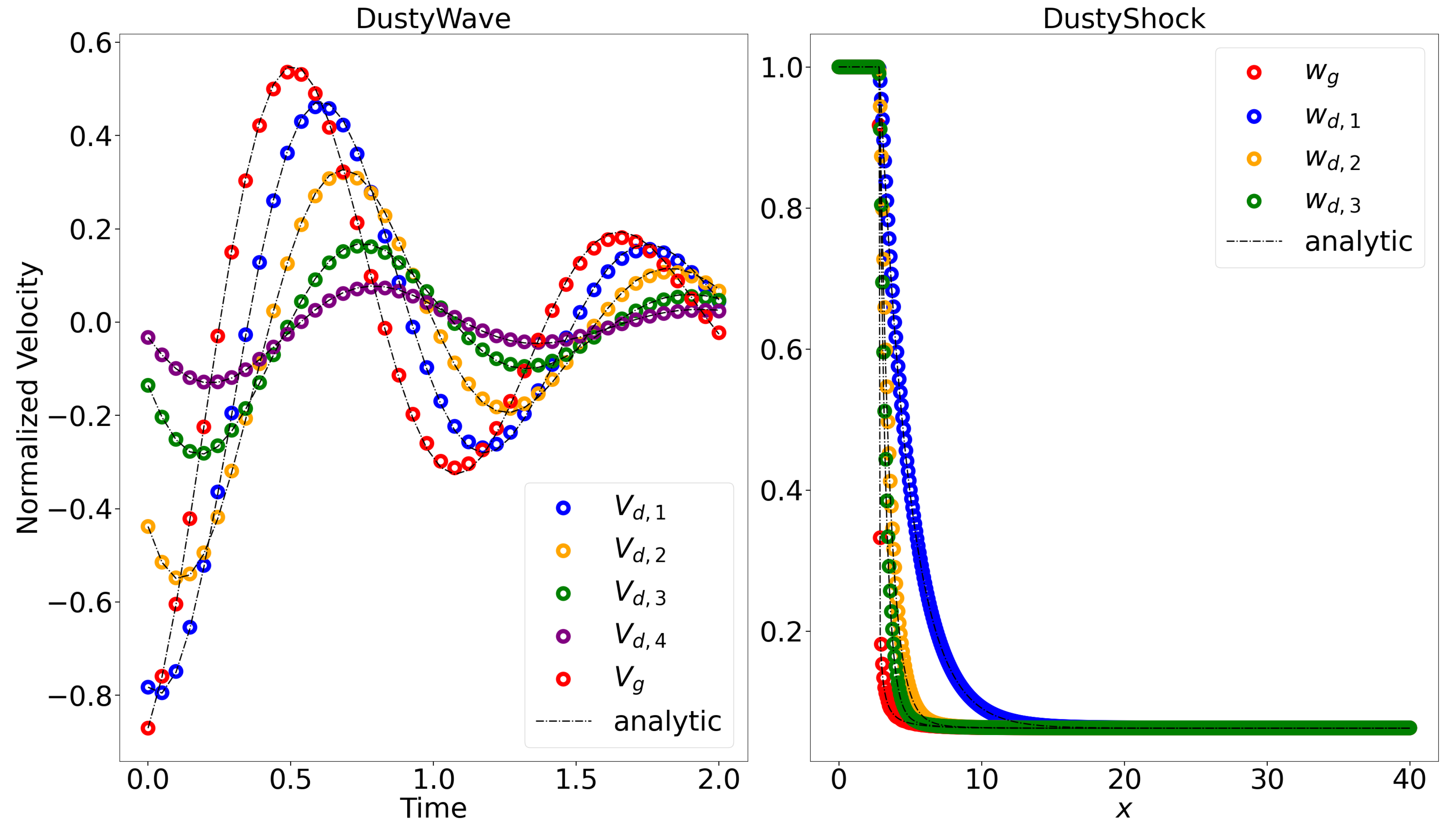
Small stopping time



Dust dominated



$$\frac{\partial \mathbf{U}}{\partial t} + \nabla \cdot \mathbf{F}(\mathbf{U}) + \mathbf{f}_{\text{src,ext}} = \mathbf{f}_{\text{drag}}$$



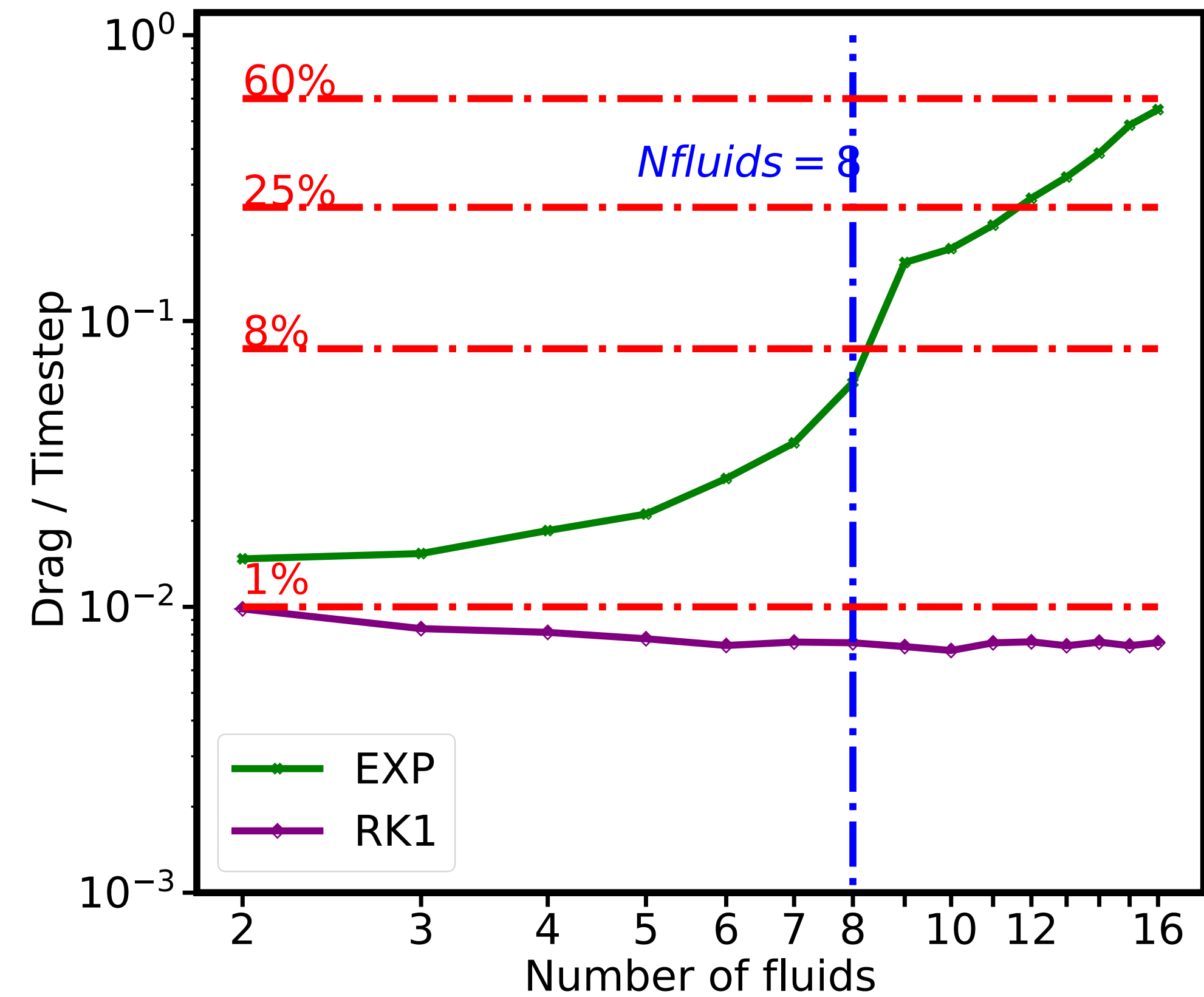
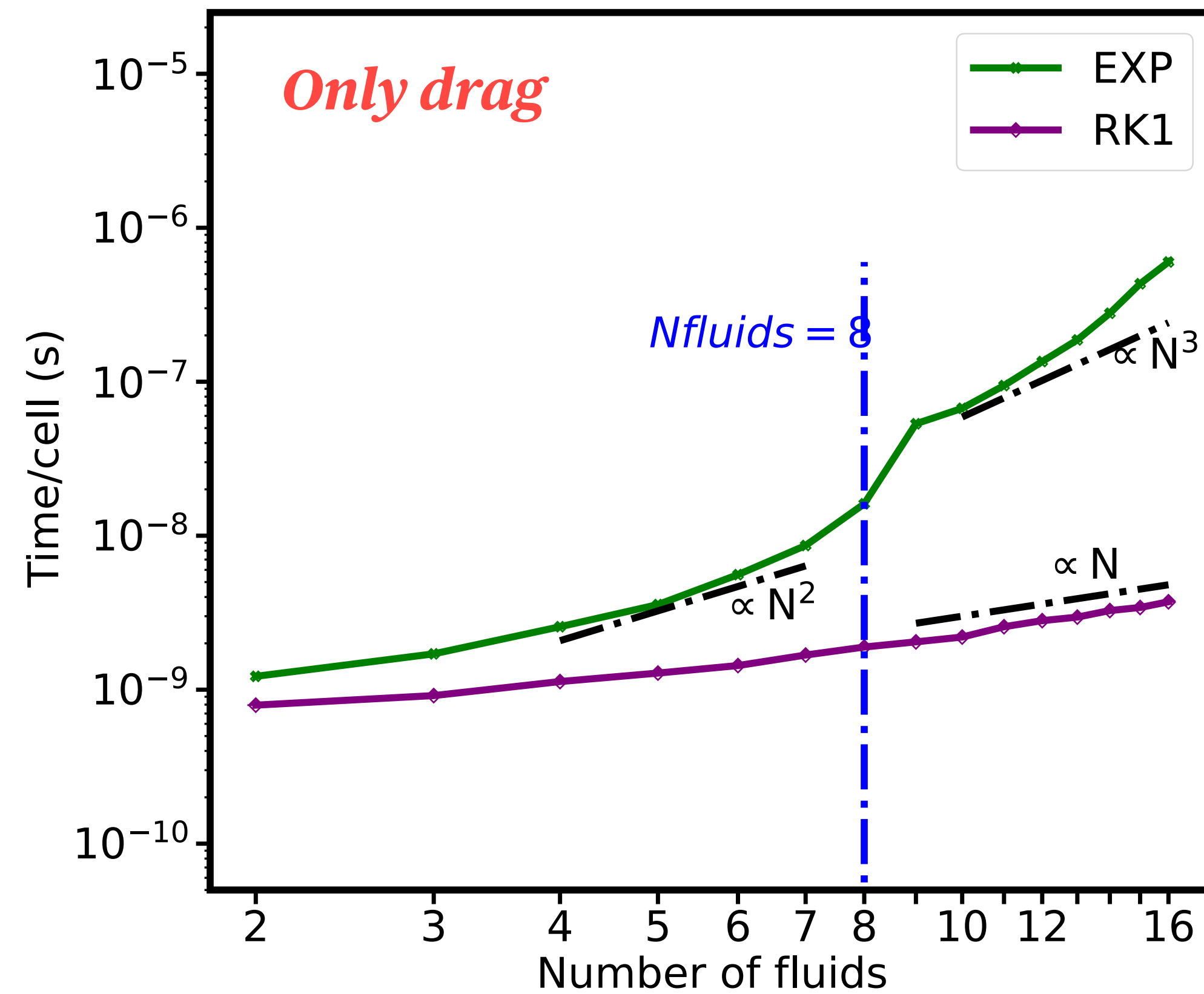
Sewanou et al, 2025

GPU A100

Dustywave test

Nsight Compute + Nsight Systems

64^3 cells



Sewanou et al, 2025

Performances on GPU (with Dyablo & Shamrock)

Static array

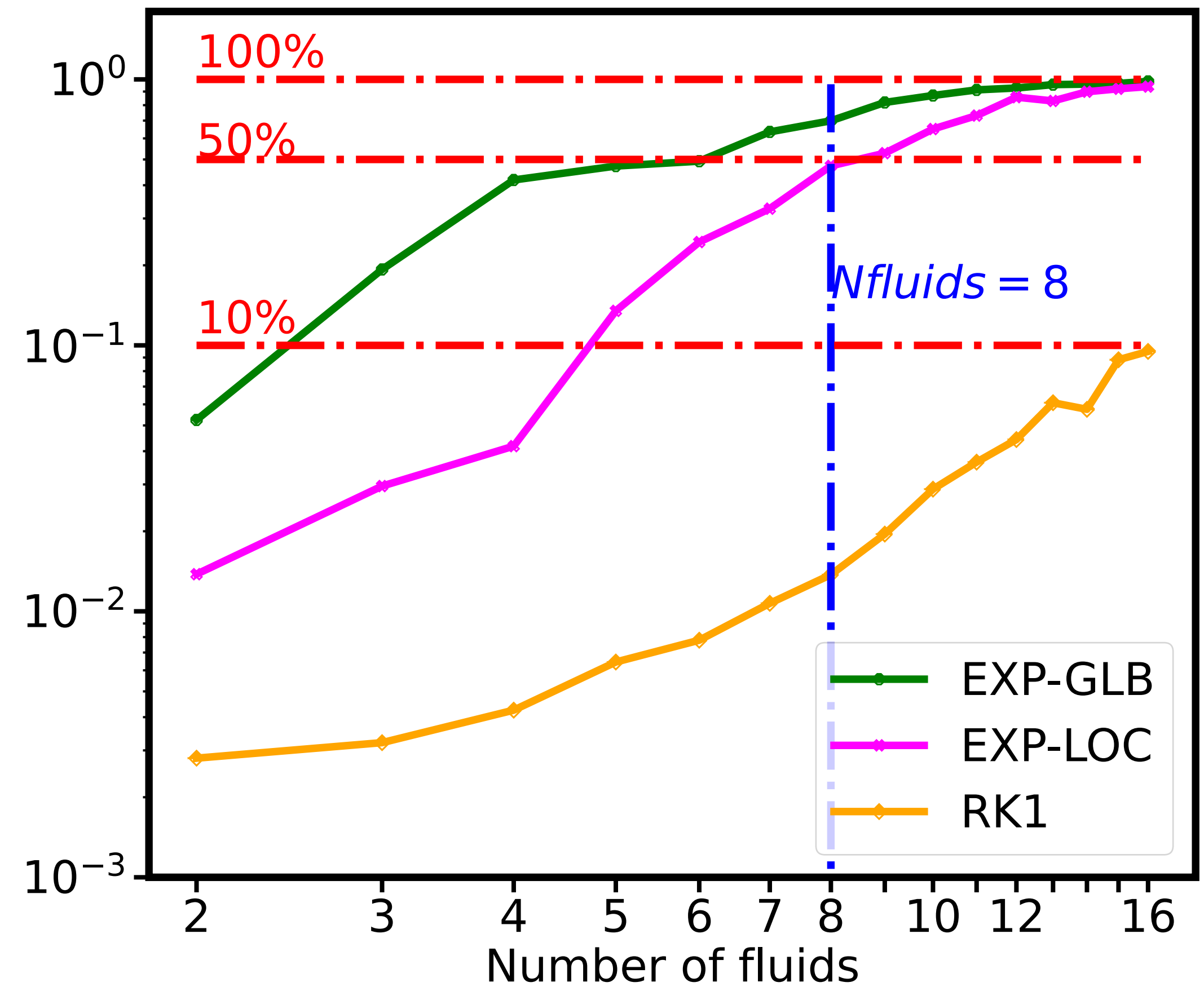
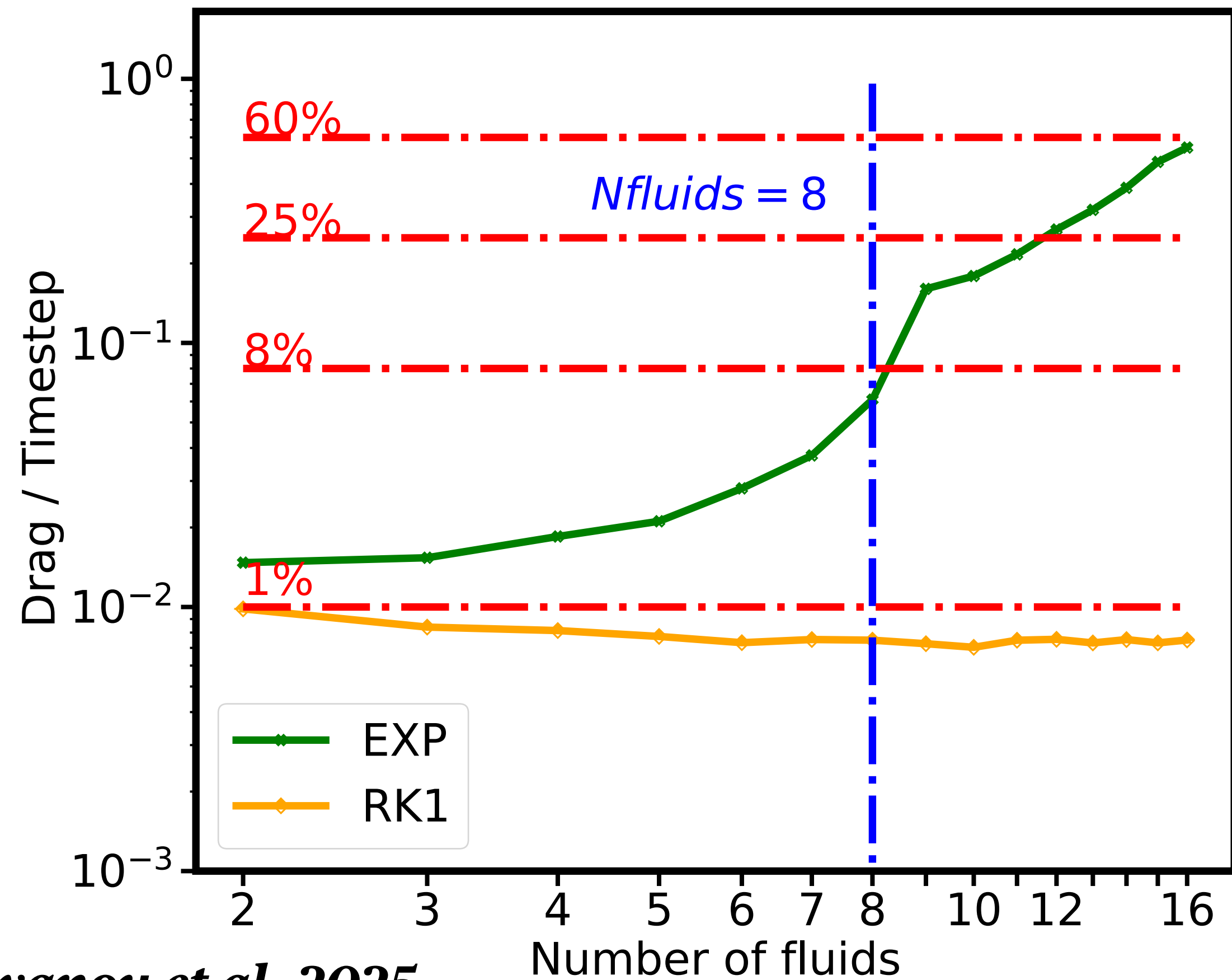
Dustywave test

Dynamic array

64^3 cells

DYABLO

SHAMROCK



Sewanou et al, 2025

High accurate drag solver on GPUs with Taylor based scaling and squaring matrix exponential computation

Time step independent

Implementation in Dyablo and Shamrock

Acceptable computational budget up to 16 dust species with current implementation

Application for other linear(ized) source terms (radiative transfer, chemistry)